

**Evolution over two decades of the tropical clouds in a subsidence area and their relations to large-scale environment**

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## Abstract

The interannual variability of cloud properties in a tropical subsidence area (South Atlantic Ocean) is examined using 23 years of ISCCP cloud fractions and optical depths, complemented with ISCCP/Meteosat visible reflectance and a four-years comparison with CALIPSO-GOCCP products. The mean seasonal cloud properties are examined in the area, as their interannual evolution. Circulation regimes (characterized with the SST and  $w_{500}$  from NCEP and ERA-Interim) that dominate summer and winter are also examined, and atmospheric situations are classified in five circulation regimes: ascending air masses, and moderate or strong subsidence with warm or cold SSTs. We examine the mean cloud cover, optical depth, and reflectance in each regime and their evolution in time over 23 years. Observational results (mean values and interannual variability) are compared with simulations from the IPSL and CNRM climate models (part of the CMIP5 experiment), using simulators to ensure that differences can be attributed to model defects.

It results that regime occurrence strongly depends on the dataset (NCEP or ERA-Interim), as do their evolution in time along 23 years. The observed cloud cover is stable in time and weakly regime-dependent, whereas the cloud optical depth and reflectance are clearly regime-dependent. Some cloud properties trends actually do exist only in some particular regimes. Compared to observations, models underestimate cloud cover and overestimate cloud optical depth and reflectance. Climate models poorly reproduce regime occurrence and their evolution in time, as well as variations in cloud properties associated with regime change. It means that errors in the simulation of clouds from climate models are firstly due to errors in the simulation of the dynamic and thermodynamic environmental conditions.

## 1. Introduction

Cloud response to anthropogenic forcing remains one of the main uncertainties for model-based estimates of climate prediction evolution [Soden *et al.* 2006; Webb *et al.* 2006; Ringer *et al.* 2006]. In the Tropics, the response of low level tropical clouds (below 440 hPa) to anthropogenic forcing is highly variable from one climate model to another, suggesting that low-level clouds contribute significantly to tropical climate cloud feedback uncertainties [Bony and Dufresne 2005]. Bony *et al.* [2004] showed that tropical low-clouds have a moderate sensitivity to temperature, but their statistical weight is so important that they could have a large influence on the tropical radiation budget. As a consequence, it is necessary to study low-level clouds in a context of global warming.

Tropical low clouds and their relations to dynamic and thermodynamic variables have been widely studied in the past: in the Pacific Ocean [Clement *et al.* 2009; Klein *et al.* 1995; Norris 1998; Norris and Klein 2000; Lau and Crane 1995; Kubar *et al.* 2010], in the Atlantic Ocean [Zhang *et al.* 2009; Mauger and Norris 2010; Oreopoulos and Davis 1993], and in all tropical oceans [Williams *et al.* 2003; Yuan *et al.* 2008; Klein and Hartman 1993; Sandu *et al.* 2010; Rozendal and Rossow 2003; Medeiros and Stevens 2009; Bony *et al.* 2004]. Only a few papers are dedicated to the interannual variability of low-clouds: Clement *et al.* [2009] used fifty years of low cloud observations in the Pacific Ocean and showed for example a positive trend of the total cloud fraction at the end of the 90's, associated with similar trends in the thermodynamical variables; Oreopoulos and Davies [1993] used cloud satellite observations in two tropical oceanic locations to study the effect of temperature variations on the cloud albedo, in particular its monthly variation during five years.

The current paper aims at characterizing the interannual variability of south-Atlantic tropical clouds located in the 0°/30°S – 30°W/8°E square (Fig. 1a), under predominance of subsidence air motion. We have chosen a larger region than the “Namibian” square used in

the KH93, RR03 and Zh09 studies (Tab. 1) because this area is a location of maximum stratus (KH93), and the goal of the current study is to examine all types of clouds associated with subsidence conditions. The high precipitation isolines in Fig. 1a show that the region under study is exposed to the ITCZ (Inter-Tropical Convergence Zone) in the northern edge in DJF (December – January – February), but not in JJA (June – July – August). Most of the time, this region is exposed to the descending air masses of the Hadley cell; in JJA its southern edge could be influenced by the subtropical anticyclone [Venegas *et al.* 1996]. This region contains both opaque stratocumulus clouds along the African coast, and shallow cumulus clouds westwards (Tab. 1, 3<sup>rd</sup> line). Like previous studies (RR03, Zh09 for the Atlantic Ocean; Klein *et al.* 1995; Norris 1998; Norris and Klein 2000 for equivalent subsidence locations in the Pacific Ocean), we analyze DJF and JJA independently because those are opposite seasons in terms of ITCZ influence.

The first objective of this paper is to analyze (i) the evolution of monthly mean cloud radiative properties over two decades in a region of subsidence, and (ii) the evolution of the concomittent dynamic and thermodynamic atmospheric properties. We try to determine if there is a robust relationship between these environmental variables (from reanalysis) and the observed cloud radiative properties (seasonal averaging and spatial resolution of  $2.5^\circ \times 2.5^\circ$ ), and if this relationship is stable over two decades. It would suggest that we could know cloud radiative properties when dynamic and thermodynamic conditions are known. Moreover, our confidence in model-based predictions of future climate, depends on the ability of models to simulate realistically the current climate. The second objective of this paper is to evaluate the ability of two climate models to reproduce the evolution over two decades of (i) the observed cloud properties, (ii) the concomittent dynamic and thermodynamic atmospheric conditions, and (iii) the relationship between these environmental conditions and cloud radiative properties.

Cloud properties are first characterized using satellite observations (Sect. 2). Then, we focus (Sect. 3) on the characterization of dynamical and thermodynamical regimes and their evolution with time. Sect. 4 describes the cloud properties associated with each regime and their interannual variability. In each section, the main results (i.e. interannual trends) are compared with simulations from the IPSL (Institut Pierre Simon Laplace 4, Hourdin et al. 2012) and CNRM (Centre National de Recherches Météorologiques, Voldoire et al. 2011) climate models. Discussion and conclusion are drawn in Sect. 5.

## 2. Cloud satellite observations

### 2.1. Satellite data

Most papers studying tropical clouds (Tab. 1) are based on satellite observations. Many of those use ISCCP (International Satellite Cloud Climatology Project, *Rossow et al.* 1991a and b; 1993; 1996; 2004) Cloud Fraction (CF), e.g: *Clement et al.* [2009], *Klein and Hartman* [1993], *Rozendal and Rossow* [2003], *Williams et al.* [2003], *Medeiros and Stevens* [2009], *Zhang et al.* [2009], *Lau and Crane* [1995], *Oreopoulos and Davies* [1993]. A few papers study tropical low clouds with A-Train observations that give more detailed information on cloud properties: *Sandu et al.* [2010] used cloud types and cloud fraction from MODIS (Moderate Resolution Imaging Spectroradiometer), associated with collocated observations of water vapor and precipitation; *Mauger and Norris* [2010] used both MODIS and CERES (Clouds and the Earth's Radiant Energy System) to study the influence of previous meteorological conditions on sub-tropical cloud properties; *Kubar et al.* [2010] used the more complex observations from CALIPSO (Cloud Aerosol Lidar and Infrared Pathfinder) and CloudSat along a tropical cross-section during one year.

Here, we use three datasets to characterize clouds: cloud fractions and optical depths from ISCCP D2 products [*Rossow et al.* 1991 and 1996; *Rossow and Schiffer* 1991], cloud

fractions from CALIPSO-GOCCP (CALIPSO – GCM Oriented CALIPSO Cloud Product, *Chepfer et al.* 2010), and visible reflectance from ISCCP/Meteosat DX products [*Desormeaux et al.* 1993].

#### *a) ISCCP cloud fraction and optical depth*

The ISCCP analyzes satellite radiance measurements to retrieve cloud fraction and optical depth. The same algorithms are applied to several spaceborne instruments, such as GOES (Geostationary Operational Environmental Satellite) instruments and Meteosat, or polar orbiters, in order to get long-term information on clouds at global scale. We used the cloud fraction for low-level clouds ( $CF_{low/ISCCP}$ , cloud top pressure  $P_{top}$  between ground and 680 hPa), mid-level clouds ( $CF_{mid/ISCCP}$ ,  $P_{top}$  between 680 hPa and 440 hPa), and high-level clouds ( $CF_{high/ISCCP}$ ,  $P_{top}$  under 440 hPa) as well as the total cloud fraction  $CF_{ISCCP}$  (sum of CF at the three pressure levels) and optical depth  $\tau$ , results of the D2 data. Because cloud fraction retrieval is based on passive remote-sensing measurements, there is no overlap between  $CF_{low/ISCCP}$ ,  $CF_{mid/ISCCP}$  and  $CF_{high/ISCCP}$ , hence  $CF_{ISCCP}$  never exceeds 100%.

We analyzed 23 years (1984 to 2006) of monthly-mean data averaged seasonally (in DJF and JJA) and spatially at a horizontal resolution of  $2.5^\circ \times 2.5^\circ$ . These observations are averaged over the diurnal cycle, unlike the observations used hereafter.

#### *b) CALIPSO-GOCCP cloud fraction and vertical profile*

The CALIPSO satellite, launched in 2006, holds the lidar CALIOP (Cloud-Aerosol Lidar with Orthogonal Polarization), which allows the characterization of the cloud vertical structure. The CALIPSO-GOCCP [*Chepfer et al.* 2010] was initially designed to evaluate cloudiness in Global Circulation Models. It is derived from CALIPSO Level 1 NASA (National Aeronautics and Space Administration) products [*Winker et al.* 2009] and contains

four types of files, including seasonal cloud fraction maps at three levels of altitude (low-, mid-, and high-level defined consistently with ISCCP). As this retrieval is based on active remote sensing, there can be an overlap between the low-level cloud fraction ( $CF_{low/GOCCP}$ ), the mid-level one ( $CF_{mid/GOCCP}$ ), and the high-level one ( $CF_{high/GOCCP}$ ); and the sum of the three can be larger than 100%. Nevertheless, the total cloud fraction  $CF_{GOCCP}$  detects if there is a cloud in the column (it does not correspond to the sum of the low, mid and high cloud fraction) and cannot exceed 100%. We also used vertical profiles of cloud fraction ( $CF_{GOCCP3D}$ ) at 40 equidistant levels (480m) from the ground to 20 km of altitude.

We analyzed four years (2007 to 2010) of CALIPSO-GOCCP seasonal mean (DJF and JJA) cloud fractions. Initially, the CALIPSO-GOCCP cloud detection is done at the full CALIOP Level 1 horizontal resolution (330 m along track and 75 m across track) to allow the detection of small-size fractionated boundary layer clouds; cloud occurrences are then statistically summarized over  $2.5^\circ \times 2.5^\circ$  grid box consistently with ISCCP cloud fraction. The CALIPSO-GOCCP cloud fraction reported here are collected in day-time (1330 LST, A-train orbit). In the current study, this product helps to understand the vertical distribution of clouds from one season to another, and from one regime to another, but is not used to draw any trend.

### *c) ISCCP/ Meteosat visible reflectance*

To consolidate our analysis, and avoid making it dependent of the inversion algorithms, we also use Meteosat Level 1 reflectance measurements (visible, wavelength near  $0.6 \mu\text{m}$ ) from ISCCP/Meteosat DX files. Reflectance can be seen as a proxy of cloud fraction combined to cloud optical depth: a radiative signature of the cloud scene.

Since we focus on a given geographical zone, the satellite viewing direction is constant; moreover the analysis is not impacted by the change of instruments as shown by

clear sky evolution in Appendix A. We use reflectance observed at a constant time 1200 UTC (3-hours time slot, as close as possible to the CALIPSO overpass time), seasonally averaged (DJF and JJA) during 23 years. The horizontal resolution is  $0.5^\circ \times 0.5^\circ$ , and the reflectance uncertainty is 0.004. Since the satellite azimuthal relative angle changes with the season, DJF and JJA reflectance cannot be compared in a quantitative way. Nevertheless, as the angles are the same every year at the same date and same location, the same seasons can be compared over different years, in order to study the interannual variability (Appendix A). Typically, stratus (St) are associated with higher reflectance than stratocumulus (Sc), cumulus (Cu) being the less reflective.

## 2.2. Cloud properties observed by satellite in the region under study: main patterns

### *a) Spatial distribution*

Reflectance value is high in three regions (Fig. 1i-j): northeast, southwest and mid-east, all with low standard deviations suggesting homogeneity of cloud solar reflectivity over years. This is consistent with previous studies that found a lot of stratus in this area, even if the area of maximum reflectance is not completely contained in the Namibian square studied previously (KH93, RR03, Zh09).

The mid-east region is characterized by important cloud fractions (about 0.8) observed by both GOCCP (Fig. 1b-c) and ISCCP (Fig. 1e-f) with low standard deviations (not shown), and large optical thickness ( $\tau \sim 8$ , Fig. 1g) with important standard deviations (not shown). This suggests that low clouds are present most of the time in this region but their cloud optical thickness is highly variable (probably due to transitions between different types of low clouds). This region corresponds with the “Namibian” area studied in KH93 who show that stratus clouds are more frequent in JJA than in DJF (0.67 versus 0.6): opposite result is found



for  $CF_{ISCCP}$  (Fig. e versus Fig. 1f) probably due to the large amount of clouds at higher levels (Fig. 1d).

The southwest region shows large cloud fraction (about 0.8) in ISCCP and GOCCP with smaller values of optical thickness ( $\tau \sim 4$ , Fig. 1i-j). Examining maps for low-/mid-/high-cloud fractions from ISCCP and GOCCP (not shown) indicates that this southwest region is mostly characterized by a mix of mid- and high-clouds.

#### *b) Mean cloud properties*

When averaged in time over 23 years and spatially over the entire area under study, cloud covers (Tab. 2) show that ISCCP detects almost only low-clouds (about 40%) whereas GOCCP detects some mid- and high-clouds (up to 20%). In complement, the mean vertical profile (Fig. 1d) shows that high-clouds are often present in this area ( $CF_{GOCCP3D} \sim 20\%$ ), with a maxima around 12 km of altitude, and these high clouds are not optically thin enough to mask the important presence of low clouds.

The two seasons exhibit similar cloud covers, but the small seasonal variation (less than 10% difference) is different in ISCCP and CALIPSO-GOCCP. The ISCCP cloud fraction  $CF_{ISCCP}$  and optical depth  $\tau$  are slightly higher in DJF than in JJA (Tab. 2). This is still the case when considering each year individually (not shown). The seasonal variation from GOCCP is opposite to ISCCP: the mean cloud cover is slightly higher in JJA than in DJF (Tab. 2, Fig. 1b-c-e-f). The discrepancy between ISCCP and GOCCP occurs only in JJA in which GOCCP observes more clouds than ISCCP at all levels of altitudes. The important result is that the difference between the cloud covers from the two datasets is less than 10% for all levels of altitude: both datasets will be used together for the rest of the interpretation.

The detailed cloud vertical distribution (Fig. 1d) changes with season: JJA shows less high clouds and more low clouds ( $CF_{GOCCP3D} \sim 0.25$ ) than DJF. The low cloud variation can

be due in part to the high cloud masking low clouds in DJF. These seasonal differences in the vertical cloud distribution confirm that the cloud types are different during each season in the area under study, and that the two seasons need to be analyzed independently in the rest of this paper (even when split into regime).

The results of climate models simulations obtained by the IPSL and CNRM models in the same region and period are given in Tab. 2. Those results have been obtained with the ISCCP [Webb and Klein 2006], CALIPSO [Chepfer et al. 2008] and reflectance simulators [Konsta et al., under review] included in COSP (CFMIP Observations Satellites Package, Bodas et al. 2011). This way, “Climate Model+Simulator” (CMS) outputs are kept consistent with satellite observations. In this region and on average over 23 years, both CMS underestimate by a factor of two the cloud cover whatever the season (Tab. 2), and overestimate significantly the reflectance (up to + 0.1 in JJA), confirming that 1) climate models do not produce enough clouds in the tropics (particularly at low levels), and 2) when models do create clouds, they are optically too thick. This result is consistent with the “too few, too bright” low-level tropical cloud problem identified in CMIP5 models (i.e. Nam et al. 2012).

### 2.3. Interannual variability of cloud properties (1984-2006)

The interannual variability of cloud fraction  $CF_{ISCCP}$ , optical depth  $\tau$  and reflectance  $ref$  is large over 23 years (Fig. 2) in both DJF and JJA. The trend of each variable, i.e. the value of the linear regression slope multiplied by the number of years is given in each subtitle in Fig. 2 (it is outlined within a rectangle when it is superior to the value of the standard deviation). In DJF, the cloud optical depth  $\tau$  and reflectance  $ref$  have increased in 23 years, despite almost no change in the cloud cover  $CF_{ISCCP}$ . In JJA, the cloud optical depth has increased of +1.2 in 23 years and reflectance remains almost stable. Anomalies can be

isolated: for example, reflectance is enhanced by +0.3 in DJF 1997 and JJA 1985, there is a strong deficit of  $\tau$  in DJF 1988. Year 1997 corresponds with an El Niño event: the important anomalies of cloud properties for this year are consistent with *Belon et al.* [2010] results: in the tropics, the fraction of interannual variability of low-cloud cover that is related to SST variability is driven by El Niño index.

This important variability in observations is well reproduced by CNRM model, whereas it is smoother with IPSL model (Fig. 2d-e-f). The two models give very different results: they do not show the same years of maximum and minimum for example. For one single model, DJF and JJA are in phase (increase or decrease the same years, maximum and minimum the same years...) whereas it is not the case in observations.

### 3. Characterization of the atmospheric circulation in the region under study

At first order, the cloud occurrences and optical depth depend on the atmospheric circulation of the air masses; as a consequence, any change in the atmospheric circulation will affect cloud properties. In order to understand if the observed interannual variability of the cloud properties (and its anomalies) is mainly due to variations of atmospheric circulation or to the change in cloud physical processes we will separate the cloud population in atmospheric circulation regimes, based on the analysis of dynamic and thermodynamic variables: in the following, it will be only called “regime”.

#### 3.1. Definition of the regimes used in this study.

Various dynamic and thermodynamic variables are used in the literature to characterize the atmospheric environment in the Tropics. *Bony et al.* [2004] showed that clouds, and in particular low clouds, are sensitive to both large-scale circulation and thermodynamic structure of the atmosphere. *Klein et al.* [1995] also concluded that the low-

level cloud fraction is better correlated to temperatures 24 to 28 hours previously, than to simultaneous ones, and that this cloud fraction is linked to atmospheric circulation at the interannual scale. The Sea Surface Temperature (SST), the Lower Tropospheric Stability (LTS) and the vertical velocity of air at 500 hPa ( $w_{500}$ ) are also frequently used to characterize the dynamic and thermodynamic state of the atmosphere. Wi03 showed that the clouds response depends more on changes in both  $w_{500}$  and SST than on changes in SST alone. *Medeiros and Stevens* [2009] suggested that  $w_{500}$  alone is not good to separate low clouds, but is useful with LTS:  $w_{500}$  identifies subsidence and LTS separates cloud types in these subsidence motions. More generally, several approaches have been followed, using only  $w_{500}$  as in *Bony et al.* [2004], only LTS as in *Wyant et al.* [2009], or a combination of both as in *Medeiros and Stevens* [2009]. In this study, we privileged the two  $w_{500}$  variable for the large-scale dynamics, and SST variable for the thermodynamics in the following, the combination of these two variables will be called “environmental variables”.

Figures 3 shows the distribution of SSTs and  $w_{500}$  over the region, as simulated by NCEP (National Centers for Environmental Prediction, Fig.3a-b) and ERA-Interim (ECMWF Re-Analyses, Fig 3.c-d) reanalyses, and IPSL (Fig. 3g-h) and CNRM (Fig. 3i-j) climate models. The seasonal variation of SST and  $w_{500}$  is roughly consistent with the position of the Hadley cell in the four models: the SST is colder in JJA than in DJF, and air masses ascending more ( $w_{500} < 0$ ) in DJF than JJA. Reanalyses show higher variability than climate models, suggesting than the latter reproduce the mean values for the SST (but not for  $w_{500}$ ), whereas the real variability is poor for models. In JJA for example, strong subsidence is too frequent and weak subsidence is missing for CNRM model, and strong subsidence never appears for IPSL model.

The PDF (Probability Density Function) of  $w_{500}$  and SST from NCEP (not shown) have been examined and values are split following sections of the curves. It leads the

separations of the dataset in five regimes (white lines in Fig.3) that are studied independently hereafter: an ascending regime (“As”) associated with deep convection, strong subsidence with cold (“SSu-cold”) or warm (“SSu-warm”) SSTs (associated with Stratocumulus), and weak subsidence with cold (“WSu-cold”) or warm (“WSu-warm”) SSTs associated with trade-wind cumulus). Weak and strong subsidence are separated at 30 hPa/day, and warm and cold SSTs are separated at 298,5 K (respectively 296,5 K) in DJF (respectively JJA).

### 3.2. Interannual variability of regime distribution (1984-2010)

The occurrence of each regime evolves in time over 27 years (1984 to 2010, in order to get both ISCCP years and GOCCP years). Figure 4 presents the interannual evolution of their occurrence as an anomaly. The interannual trends (calculated as in Sect. 2.3) that are statistically significant (underlined by a rectangle in the subtitle of each subplot in Fig.4), represent about half (11 out of the 20) of the values reported here. A single regime in one season exhibits a consistent interannual trend statistically robust in both datasets (NCEP and ERA-Interim): the occurrence of the As regime decreases by 12-15% in 27 years in DJF. For all other regimes, the trends obtained with the two reanalyses are inconsistent and/or not statistically representative. Moreover, based in NCEP reanalyses, the As regime is dominant in DJF from 1984 to 1990 whereas SSu-cold dominates 1998 to 2010 (not shown).

Similarly to the reanalyses, climate models (Fig. 4f and g) do not suggest a significant change in regime occurrence with time, except for the WSu-warm regime in DJF (only for CNRM model). Nevertheless, this regime becomes more frequent with time in DJF for CNRM model, which is in contradiction with the trend produced by ERA-Interim and NCEP reanalyses. Moreover, trends that are significant in the reanalyses are not significant in the models.

This suggests that both climate models are far of reproducing the occurrence of regimes given by the reanalyses, but they are even more far away of reproducing their evolution in time. In particular, the models predict that ascending branch of the Hadley cell (As regime) is more frequent, which is not the case in the reanalyses.

## **4. Analysis of cloud properties for each regime**

### 4.1. Characterization of cloud properties in regimes

To assess how robust is the mean cloud properties dependence on the regime, results obtained with the different satellite datasets (ISCCP, CALIPSO-GOCCP) and reanalyses datasets (NCEP and ERA-Interim) are reported in Fig. 5.

Observations (blue and green) are consistent for both datasets: the mean cloud fraction varies slightly (between 0.4 and 0.6) when the regime changes, whereas the mean optical depth (between 3 and 6) and the mean reflectance (between 0.15 and 0.25) are significantly regime dependent, and seasonally dependent. In DJF, the larger optical depth and reflectance are associated with strong subsidence regime (stratocumulus), which optical depth decreases significantly in winter. In JJA, the cloud cover is about the same for all regimes, but larger optical depths (and reflectance) are encountered in deep convection (As regime). These results show that regimes do not drive significantly the mean cloud cover in the region, but do drive the mean cloud optical depth and reflectance, and significantly the vertical structure (Fig. 6). In particular, for the subsidence regime, the SST impacts more the vertical structure than  $w_{500}$ , consistently with Wi03 results, in particular in winter (JJA). However, a regime (as defined in this study) does not by itself completely determine the cloud optical depth and reflectance: for a given regime the mean optical depth and reflectance also depend on the season. This seasonal dependency can be explained by the change in the cloud vertical distribution as shown in Fig. 6.

Compared to the observations, CMS underestimate the cloud cover (Fig. 5a and d) by a factor of two (or more) in all regimes and seasons, except in the As regime that is better described by one CMS (CNRM). Differences between observations and CMS cloud cover can be more than a factor of three for some regimes: SSu-cold and -warm (but they are not significant in term of population, Fig.3) and WSu-cold in DJF, and SSu-warm in JJA (i.e. 41% of the population for CNRM model, Fig. 3h). It confirms that the boundary layer cloud scheme, that drives the amount of cloud forms in subsidence conditions, remains a challenging task for those two climate models.

The IPSL model errors on the reflectance (overestimate, Fig. 5c-f) and on the cloud cover (underestimate, Fig. 5a-c) likely compensate to produce correct shortwave fluxes at the top of the atmosphere (in both seasons) as already mentioned in previous section. Figure 5 shows that this error compensation applies to all regimes that are significant in term of population.

#### 4.2. Interannual variability of cloud properties using regime classification

Figure 7 shows the trends of observed cloud variables (blue and green bars) over 23 years in each regime. It means for example that  $ref$  has an increase of +0.03 in 23 years for the As regime in DJF (based on NCEP reanalyses). The cloud fraction is stable in time in all regimes (contrary to Clement et al. 2009), except in the warm strong subsidence regime where it decreases (Fig. 7a) consistently with the reflectance (Fig. 7c). In the other four regimes, cloud optical depth (Fig. 7b) and reflectance (Fig. 7c) have increased very slightly in summer (DJF) over 23 years (about +0.5 for  $\tau$  and about +0.035 for  $ref$ , in 23 years). A more important increase of cloud optical depth occurs in winter (JJA) in all regimes (about +1.5), but it is not associated with change in the reflectance that remains stable in time.

A trend is robust if it is observed in both sets of reanalyses, hence the final results are:  
(1) a decrease of cloud fraction for the SSu-warm regime in DJF, (2) an increase of optical depth for weak subsidence in DJF and for all regimes (except ascent) in JJA, and (3) an increase of reflectance in the ascent regime in DJF.

In most of the cases, the IPSL CMS does not show any robust trend in the cloud cover and reflectance. When it shows some trends (vertical arrows in Fig. 7), those are sometimes in contradiction with the observations: in the WSu-warm regime, the modelled *CF* and *ref* in JJA increase in time along the last 23 years (Fig. 7c) which is not consistent with the observations. This increase of *ref* suggests that the clouds of this specific regime reflect more solar light now than 23 years ago. But the observations disagree with this modelled reflectance trend.

## 5. Conclusion

We have examined cloud properties in a tropical subsidence area (south Atlantic Ocean) using 23 years of ISCCP cloud fractions and optical depths, complemented with 23 years of visible reflectance from ISCCP/Meteosat, and cloud vertical profiles from CALIPSO-GOCCP collected during four years. We first studied the mean cloud properties (cloud cover, optical depth, and reflectance) in DJF and JJA. The region under study contains about 40% of low-level clouds and 20% of high-clouds (around 12 km). The difference between ISCCP and GOCCP cloud cover is less than 10%, but the small seasonal variation is not consistent between the two dataset. Then we looked at the interannual variability of cloud properties over 23 years using ISCCP: cloud cover is stable in time and cloud optical depth exhibits a positive trend (+0.55 in DJF and +1.2 in JJA).

We compared the observational results with output from climate models (IPSL and CNRM models within CMIP5), coupled with COSP to ensure that differences can be



389 attributed to model defects. Models underestimate cloud cover by a factor of two, and  
390 overestimate reflectance (+0.1). The CNRM model produces a stable cloud cover in time, in  
391 agreement with observations, whereas the IPSL model shows a significant and unrealistic  
392 positive cloud cover trend over the years.

393         As the cloud formation and properties are primarily driven by atmospheric dynamic  
394 and thermodynamic variables, we examined the regimes (characterized with the SST and  $w_{500}$   
395 from NCEP and ERA-Interim) that dominate DJF and JJA. We classified atmospheric  
396 situations in five categories: ascending air masses, moderate subsidence with warm / cold  
397 SSTs, strong subsidence with warm / cold SSTs. The occurrence of each regime in the region  
398 depends significantly on the dataset used (NCEP or ERA-Interim). The evolution in time of  
399 the occurrence of each regime along the 23 years is different and inconsistent in both datasets,  
400 except for the “ascending air” regime: its occurrence decreases significantly in time (of more  
401 than 10%) according to both datasets. The occurrence of all regimes is poorly reproduced by  
402 both climate models (CNRM and IPSL). Moreover, both report an increase in occurrence of  
403 the “ascending air” regime contrarily to the observations.

404         We examined the relationship between environmental variables and the observed  
405 cloud properties (seasonal averaging and spatial resolution of  $2.5^\circ \times 2.5^\circ$ ). Observations  
406 indicate that the cloud cover (0.4 to 0.6) is slightly regime dependent whereas the optical  
407 depth (4 to 6) and the reflectance (0.15 to 0.25) are more significantly regime dependent.  
408 Differences between modeled and observed cloud cover and reflectance are not regime  
409 dependent.

410         We study the evolution of the relationship between the environmental variables and  
411 cloud radiative properties over two decades. The observations exhibit two robust trends over  
412 23 years in specific regimes. The optical depth increases only in weak subsidence conditions  
413 in DJF (+0.6), and for weak and strong subsidence regimes in JJA (+1). The reflectance

increases only for the ascent regime in DJF (+0.03). This later trend is reproduced by IPSL model with a smaller amplitude (+0.01). Trends detected in cloud properties before the regime separation are now explained in some regime, particularly in DJF: the decrease of cloud fraction over 23 years is explained by only one regime (strong subsidence with warm SSTs), as for the optical depth increase which is detected only for weak subsidence, whereas the reflectance increase is not detected in the subsidence (only in ascent).

In summary, this study suggests that the main difficulty to built reliable relationships between environmental variables and clouds comes from the significant uncertainties in these environmental variables produced by the different reanalyses and by climate models. It limits the ability to detect robust regime-dependent trend in the observations, and it may be the first-order limitation for models to reproduce observed clouds.

Future work will consist in extending this study to the entire tropical belt including all CMIP5 models and the same two sets of reanalyses. It will aim at determining if, at this scale, some of the regimes (and related trends) are better reproduced than others and in these cases if the link between cloud properties and environmental variable (and related trends) is better predicted by models.

## Appendix A

There are two well-known problems for the retrieval of cloud fraction using satellite passive remote sensing, in particular from the ISCCP program: the variations of the satellite angles, and the calibration of satellites when the instruments are changed. In this Annex we investigate if these problems affect the dataset used in our study.

Figure A1.a shows an example of the satellite viewing angle  $\theta_v$  for the complete area of study, for one day of the database. The values of this angle are between a few degrees and approximately  $40^\circ$ , depending on the location. Figure A1.b shows the percentage of days when  $\theta_v$  is lower or smaller than its median value by more than  $2^\circ$ , for each pixel during the time period of the study. This percentage is always lower than 4%, and Fig. A1.c shows that the concerned  $\theta_v$  do not deviate from the median by more than  $3^\circ$ . This shows that, during the 23 years of the study, the variations of  $\theta_v$  are so small that they should not be a problem for our study of reflectance trends.

Figure A2.a is the same as Fig. A1.a but for the satellite relative azimuthal angle  $\phi$ . The values of this angle are between  $0^\circ$  and  $180^\circ$ , depending on location and time. Figure A2.b is approximately the same as Fig. A1.b but for each pixel, the percentage is calculated every year for the same day, so from 23 values, in order to remove the natural variations of  $\phi$  and only consider the variations due to technical problems. Another difference is that the percentage is calculated when  $\phi$  is lower or smaller than its median value by more than  $5^\circ$ . This percentage is about 4% or 8% (one or two cases on twenty three) and Fig. A2.c shows that the concerning  $\phi$  can be different from the median value by  $50^\circ$ . Figure A3 shows the value of the solar angle for the area under study, in January (Fig. A3a) and in July (Fig. A3b). Using extreme values of these three angles (Fig. A1-A3), the correspondence between reflectance and optical depth values has been calculated. The calculation is done using a doubling/adding radiative transfer code [De Haan *et al.* 1987], assuming the atmosphere is

plane parallel infinite. The atmosphere contains a cloud composed of liquid water spherical particles of 6- $\mu\text{m}$  radius (Mie Theory). Six values of cloud optical depth ( $\tau_{calc} = 0, 1, 5, 10, 50, 100$ ) are considered and four different geometries (two extremes of January, and two extremes of July). Figure A4 is an illustration of this calculation (using a linear interpolation), and it shows that for one given optical depth, the reflectance variability is very small from one geometry to another (less than 0.1).

Figure A5 shows the variation in time of the clear sky reflectance for the complete time period by selecting, for each day, the smallest reflectance in the area. The figure shows that instrument changes are not associated with any gap in clear sky reflectance values. It follows that instrument changes are not either a problem for the current study that focuses on reflectance.

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477

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## List of Tables

**Table 1:** Review of previous studies concerning tropical low clouds and their relations to dynamic and thermodynamic variables, when the area of study includes part of total of the location of the following paper, called A. \*SON is for September to November, MAM is for March to May, MJJAS is for May to September, NDJFM is for November to March; \*\*Cloud Top Pressure; \*\*\*Sc for Stratocumulus, Cu for Cumulus.

**Table 2:** Mean values ( $0^{\circ}$  -  $30^{\circ}$ S,  $30^{\circ}$ W –  $8^{\circ}$ E) of the entire database for ISCCP (1984 – 2006) and CALIPSO-GOCCP (2007 – 2010). Equivalent simulated values from the IPSL and CNRM models have been added when available.

Reference	Important results in the A area	Location
<i>Klein and Hartman</i> 1993 [KH93]	<ul style="list-style-type: none"> <li>- SON* is season of maximum stratus &amp; maximum LTS</li> <li>- A is an area of maximum stratus</li> <li>- interannual variability in stratus are related to changes in LTS</li> </ul>	Area contained in A
<i>Rozendal and Rossow</i> 2003 [RR03]	<ul style="list-style-type: none"> <li>- <math>CF_{low}(MJJAS^*) &gt; CF_{low}(NDJFM^*)</math></li> <li>- notable differences between the low cloud areas in Pacific and the low cloud areas in Atlantic (CF, <math>\tau</math>, CTP**...)</li> <li>- more the subsidence is important, more the cloud top is low</li> </ul>	Area contained in A
<i>Sandu et al.</i> 2010 [Sa10]	<ul style="list-style-type: none"> <li>- transition of decrease CF associated with strong increase of SST &amp; decrease of LTS, &amp; free troposphere gradual humidification</li> <li>- Sc to Cu*** transition is stable from one Ocean to another, but Sc CF is higher in South Hemisphere Oceans</li> </ul>	All Atlantic Ocean
<i>Williams et al.</i> 2003 [Wi03]	<ul style="list-style-type: none"> <li>- cloud response depends more on <math>w_{500}</math> &amp; SST changes than on SST changes only</li> <li>- low clouds: many with medium <math>\tau</math>, CF more important for strong subsidence and cold SSTs</li> </ul>	All tropical oceans
<i>Medeiros and Stevens</i> 2009 [MS09]	<ul style="list-style-type: none"> <li>- <math>CF_{low}</math> increase as LTS increase but is independent of <math>w_{500}</math></li> <li>- the peak of <math>CF_{low}</math> is about 30%, very large and very low values of <math>CF_{low}</math> are rare</li> <li>- in A: a few shallow-Cu at high level, a lot of Sc at low level</li> </ul>	All tropical oceans
<i>Zhang et al.</i> 2009 [Zh09]	<ul style="list-style-type: none"> <li>- <math>CF_{low}</math> increases linearly as a function of LTS</li> <li>- <math>CF_{low}</math> &amp; LTS: both are maximum in SON* and minimum in MAM*</li> </ul>	Area contained in A
<i>Oreopoulos and Davies</i> 1993 [OD93]	<ul style="list-style-type: none"> <li>- SST has negative correlation with Albedo and CF, also for interannual variations</li> </ul>	Area contained in A
<i>Bony et al.</i> 2004 [Bo04]	<ul style="list-style-type: none"> <li>- low clouds have moderate sensitivity to temperature change but have an important statistical weight, so a large influence on the tropical Radiative budget</li> <li>- cloud Radiative forcing is high for ascendance and small for subsidence</li> </ul>	All tropical oceans

589 *Table 1: Review of previous studies concerning tropical low clouds and their relations to*  
590 *dynamic and thermodynamic variables, when the area of study includes part of total of the*  
591 *location of the following paper, called A. \*SON is for September to November, MAM is for*  
592 *March to May, MJJAS is for May to September, NDJFM is for November to March; \*\*Cloud*  
593 *Top Pressure; \*\*\*Sc for Stratocumulus, Cu for Cumulus.*

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	$CF_{ISCCP}$		$\tau_{ISCCP}$		$CF_{GOCCP}$		$ref$	
	DJF	JJA	DJF	JJA	DJF	JJA	DJF	JJA
Low – obs	0.46	0.39	–	–	0.39	0.47	–	–
Mid – obs	0.05	0.07	–	–	0.13	0.11	–	–
High – obs	0.01	0.03	–	–	0.20	0.16	–	–
Tot – obs	0.52	0.49	4.21	3.97	0.52	0.56	0.18	0.15
Tot – mod IPSL	0.18	0.17	–	–	0.21	0.20	0.2	0.3
Tot – mod CNRM	0.22	0.21	–	–	0.26	0.26	–	–

596

597 *Table 2: Mean values ( $0^{\circ}$  -  $30^{\circ}$ S,  $30^{\circ}$ W –  $8^{\circ}$ E) of the entire database for ISCCP (1984 – 2006)*  
598 *and CALIPSO-GOCCP (2007 – 2010). Equivalent simulated values from the IPSL and*  
599 *CNRM models have been added when available.*

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## List of figures

**Figure 1:** Area under study. a) Isolines of NCEP precipitable water for entire atmosphere (1984 – 200WSu-cold6), every 3 kg/m<sup>2</sup>, only larger than 35 kg/m<sup>2</sup> highlighting ITCZ area. Red is July, blue is January. Dotted square is the area of current study, dashed one is the area of KH93 study; b-c) Mean  $CF_{GOCCP}$  (2007 – 2010) (DJF-JJA); d)  $CF_{GOCCP3D}$  mean vertical profile (2007 – 2010); e-f) same as b-c) for  $CF_{ISCCP}$  (1984 – 2006); g-h) for  $\tau$ , i-j), for *ref*.

**Figure 2:** Evolution of the mean cloud properties anomaly from 1984 to 2006. a)  $CF_{ISCCP}$ ; b)  $\tau$ , c) *ref*; d)  $CF_{ISCCP}$  from models; e)  $CF_{GOCCP}$  from models; e) *ref* from models. Blue for DJF observed, red for JJA observed, black for JJA IPSL model, brown for JJA CNRM model, magenta for DJF IPSL model, green for DJF CNRM model. Horizontal dashed lines correspond to the values of standard deviations. Numbers in the titles are values of trends in 23 years, and the significant ones (i.e. superior to the variability – standard deviation) are in rectangles. Blue vertical lines indicate particular years in DJF, red ones the same in JJA, and black ones particular years in both DJF and JJA.

**Figure 3:** Log of percentage of occurrence of SST values (by classes of 1 K) versus  $w_{500}$  values (by classes of 10 hPa/day) from 1984 to 2010. a) NCEP in DJF; (b) NCEP in JJA; (c) ERA-Interim in DJF; d) ERA-Interim in JJA; e) IPSL model in DJF; f) IPSL model in JJA; g) CNRM model in DJF; h) CNRM model in JJA. (white horizontal and vertical lines indicate the limits of the five dynamical regimes and the number in white are the percentage of pixels of each regime).

**Figure 4:** a-e. Anomaly of the percentage of pixels of each dynamical regime from 1984 to 2010 (DJF in blue, JJA in red) defined by NCEP  $w_{500}$  and SST values. a) As, b) WSu-cold, c) WSu-warm, d) SSu-cold, e) SSu-warm. The horizontal dashed lines correspond to the values of standard deviations. Values of the trends in 27 years are indicated in the titles for

both NCEP and ERA-Interim reanalysis, and the significant ones are in rectangles. Vertical lines are the same as in Fig. 2. f-g. Equivalent trends (the value of the linear regression slope of the curve, multiplied by the number of years) values over 27 years for DJF (f.) and JJA (g.) calculated from NCEP, ERA-Interim, IPSL model, CNRM model. Significant trends are indicated by an arrow.

**Figure 5:** Mean values of cloud properties for the five dynamical regimes. Three first lines are  $CF$ ,  $\tau$ , and  $ref$  in DJF, three last lines the same in JJA. Blue bars are based on NCEP reanalyses, green ones on ERA-Interim reanalyses. Same results from models (IPSL and CNRM) have been added (not colour bars).

**Figure 6:** mean vertical profile of  $CF_{GOCCP3D}$  (2007 – 2010). The complete database are represented by the black dotted lines, it is then separated onto the five regimes. a) in DJF; b) in JJA. X axis is in logarithmic scale.

**Figure 7:** Same as Fig. 5 but for trend values (over 23 years) instead of mean values. A “trend” is the value of the linear regression slope of the curve, multiplied by the number of years.  $CF_{GOCCP}$  values are not drawn as there are too few years for trend estimation. Black arrows are added when the trend is superior to the variability (i.e. the standard deviation). Same results from models (IPSL and CNRM) have been added (not colour bars).

**Figure A.1:** Satellite viewing angle  $\theta_v$  (a) example of values in January, (b) percentage of values that are more than  $2^\circ$  lower or larger than the median value for the entire time period, (c) mean value of the difference between  $\theta_v$  and its median when the percentage of (b) is non-zero.

**Figure A.2:** Satellite azimuthal relative angle  $\phi$  (a) example of values in January, (b) percentage of values that are more than  $5^\circ$  lower or larger than the median value for the entire time period, (c) mean value of the difference between  $\phi$  and its median when the percentage of (b) is non-zero.

**Figure A.3:** (a) Solar zenithal angle at 1200 UTC on January 15<sup>th</sup>; (b) same as (a) but on July 15<sup>th</sup>.

**Figure A.4:** Simulation of the reflectance as a function of the optical thickness, for four representation of the satellite geometry: solar zenithal angle of 20°/angle of viewing direction with nadir of 15°/relative viewing azimuth angle of 50° (characteristic of January, red plane line), 32°/15°/50° (characteristic of January, blue plane line); 22°/42°/130° (characteristic of July, red dashed line); 50°/42°/130° (characteristic of July, red dashed line).

**Figure A.5:** Time evolution of the clear sky reflectance, during the complete period (in blue). Satellite changes are indicated by vertical red lines.

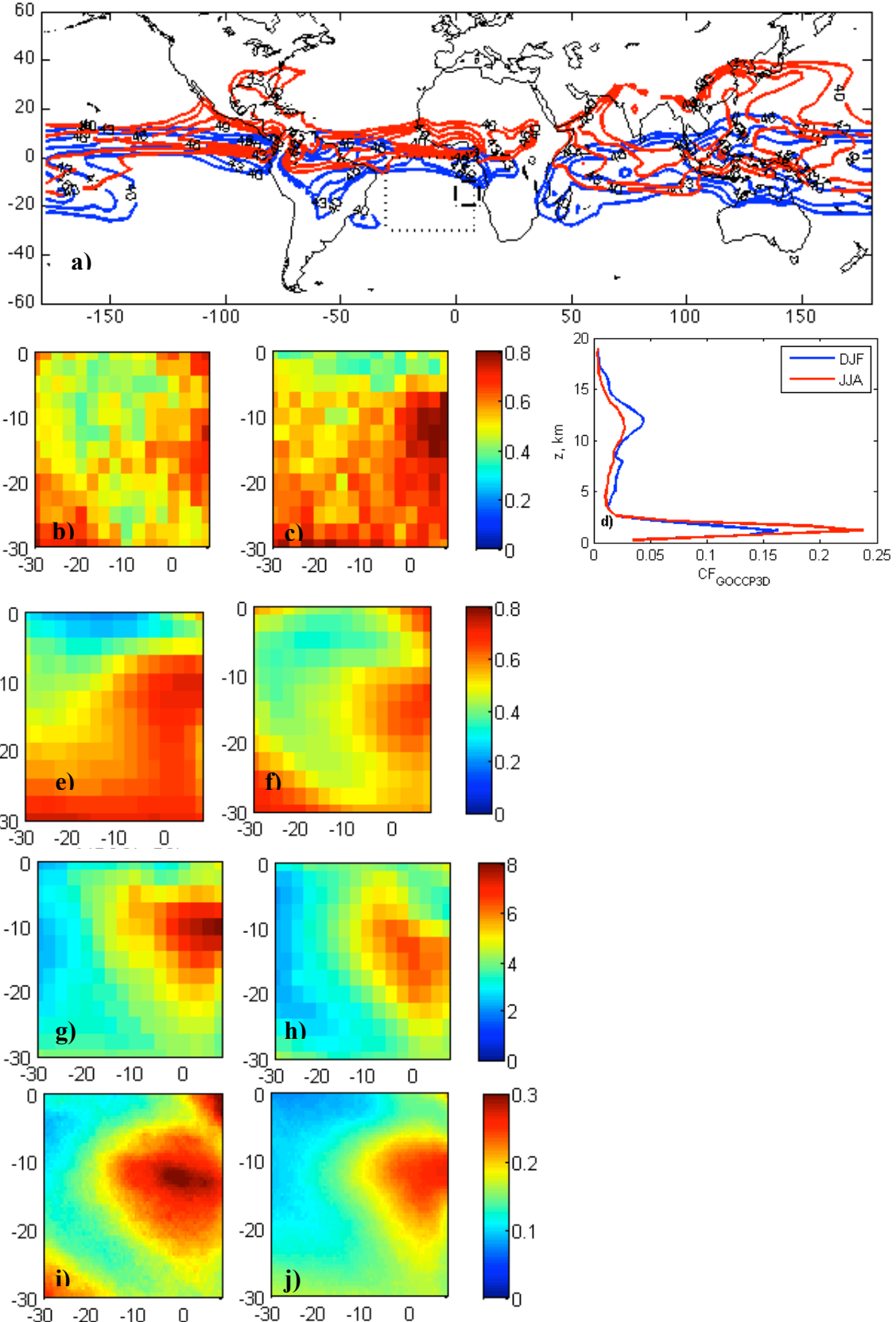


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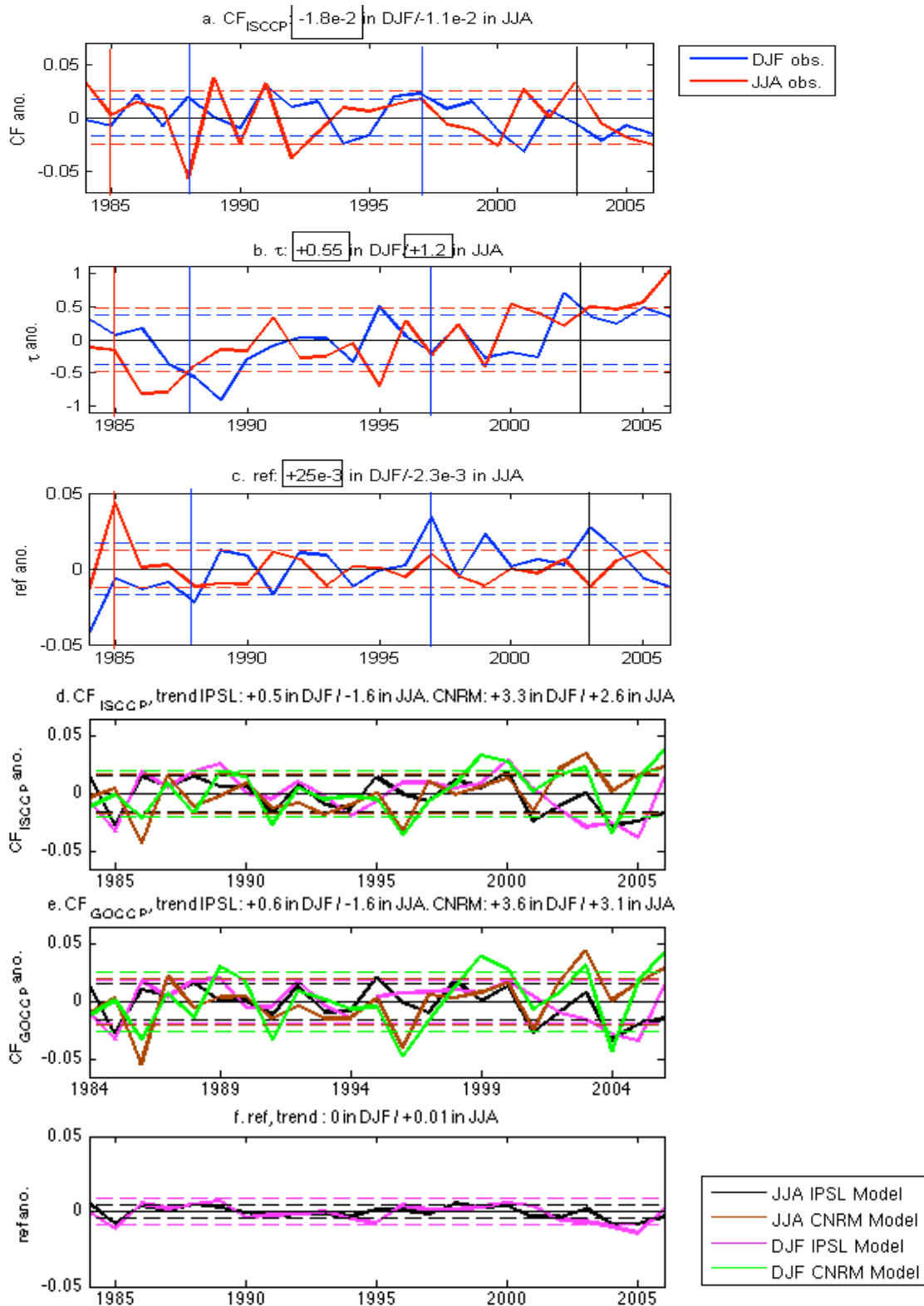


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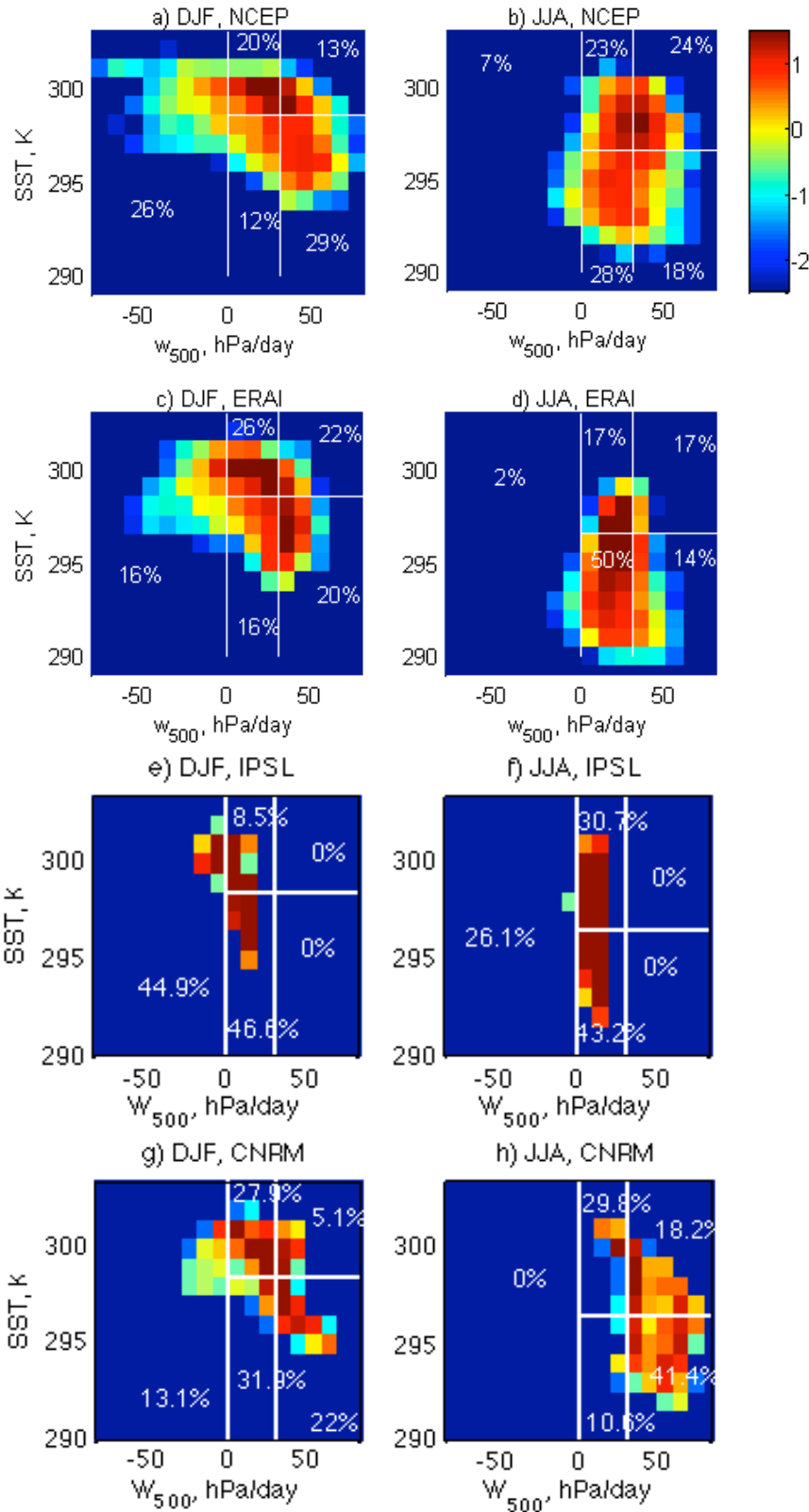


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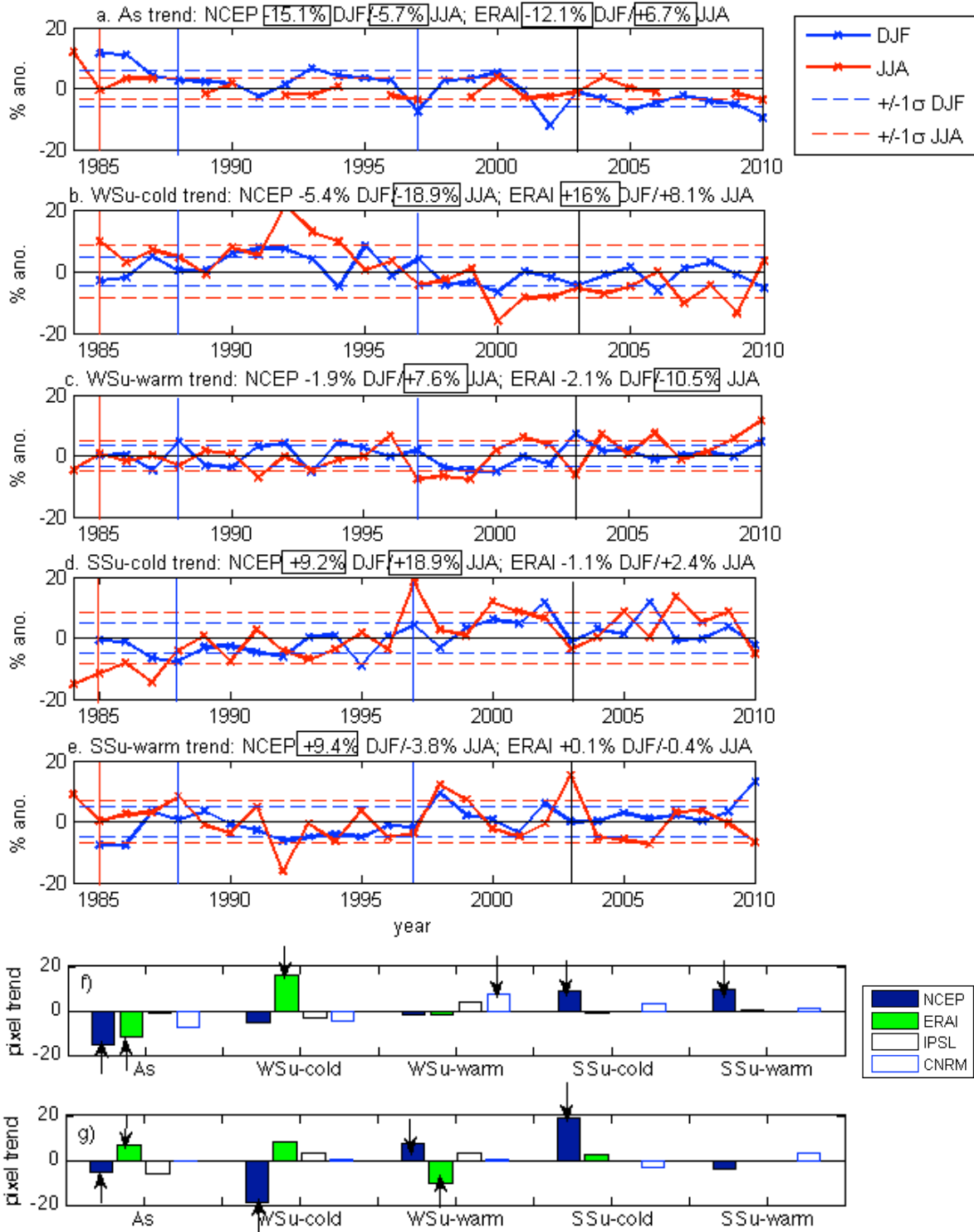


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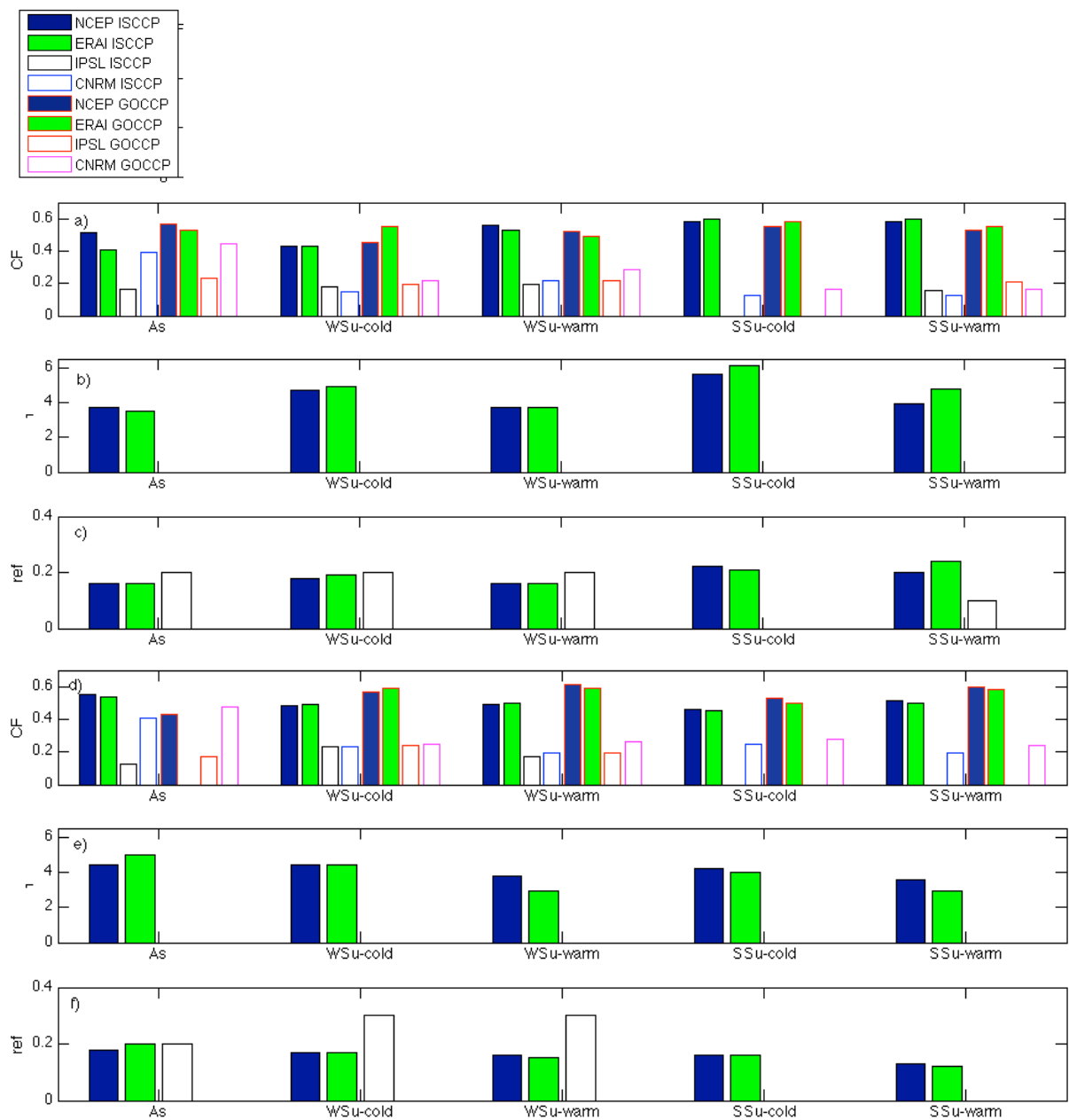


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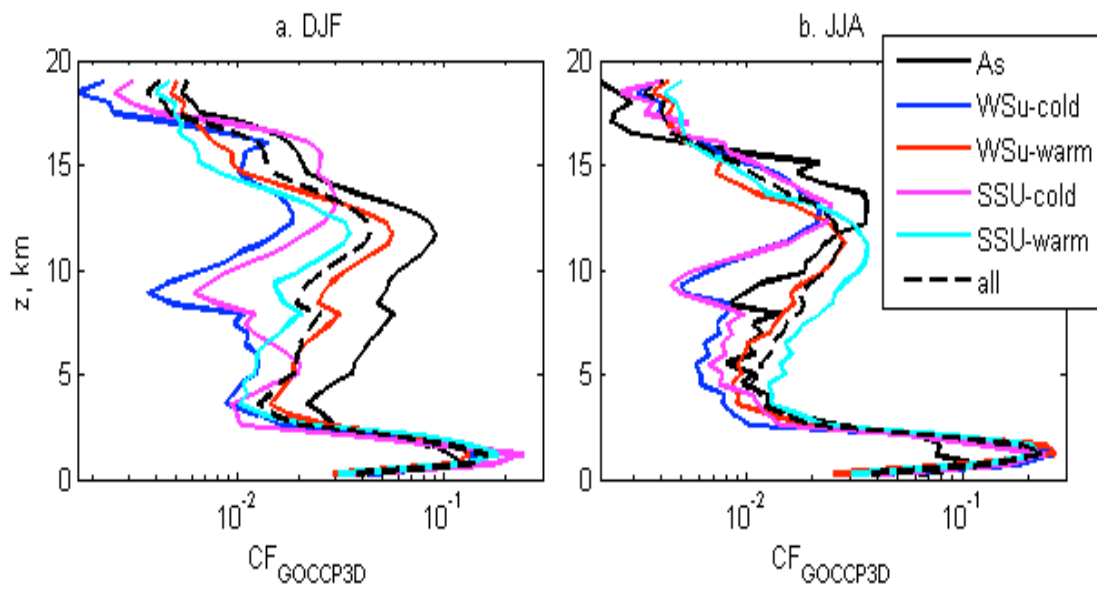


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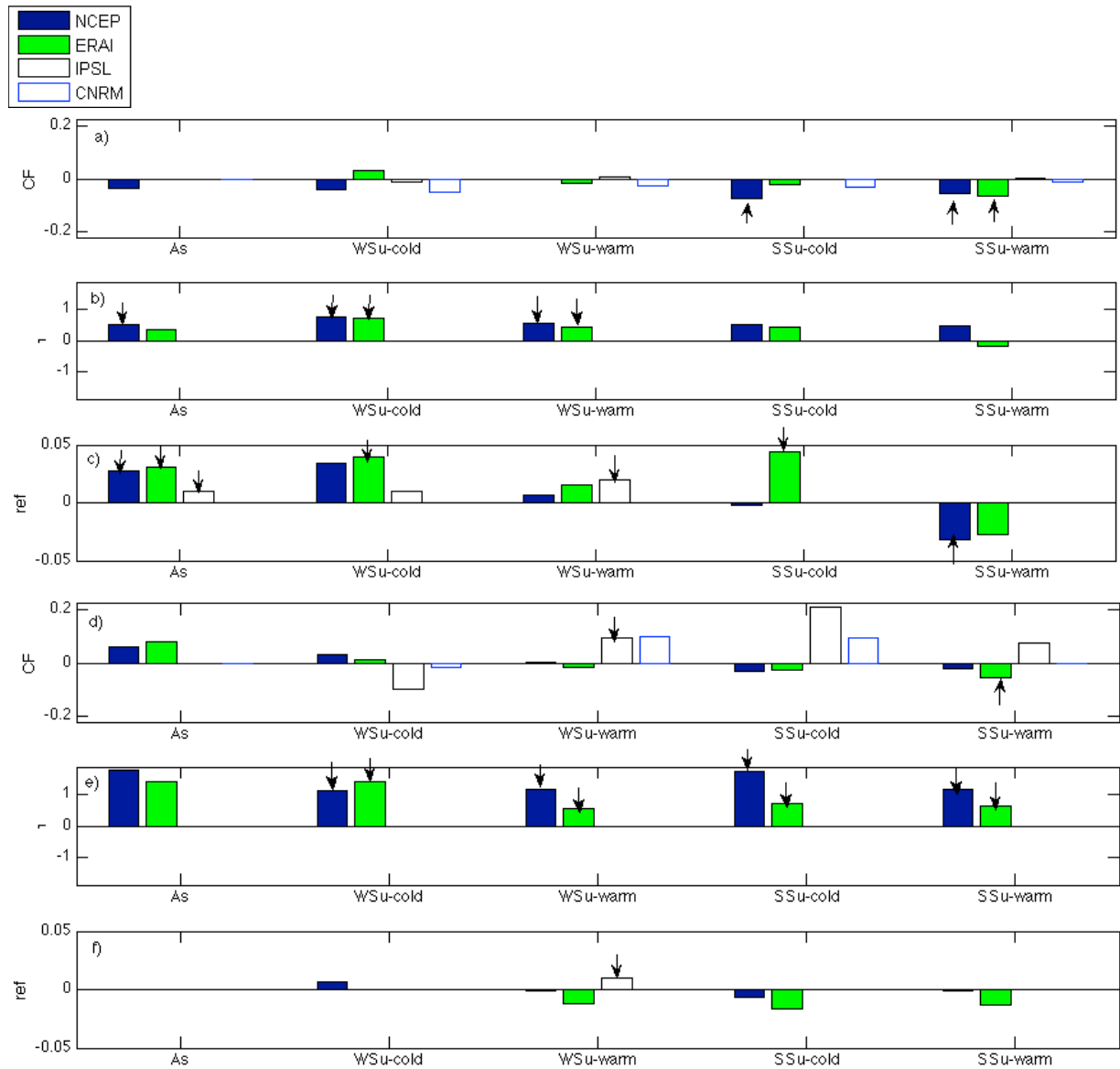


Figure 7: Same as Fig. 5 but for trend values (over 23 years) instead of mean values. A “trend” is the value of the linear regression slope of the curve, multiplied by the number of years.  $CF_{GOCCP}$  values are not drawn as there are too few years for trend estimation. Black arrows are added when the trend is superior to the variability (i.e. the standard deviation). Same results from models (IPSL and CNRM) have been added (not colour bars).

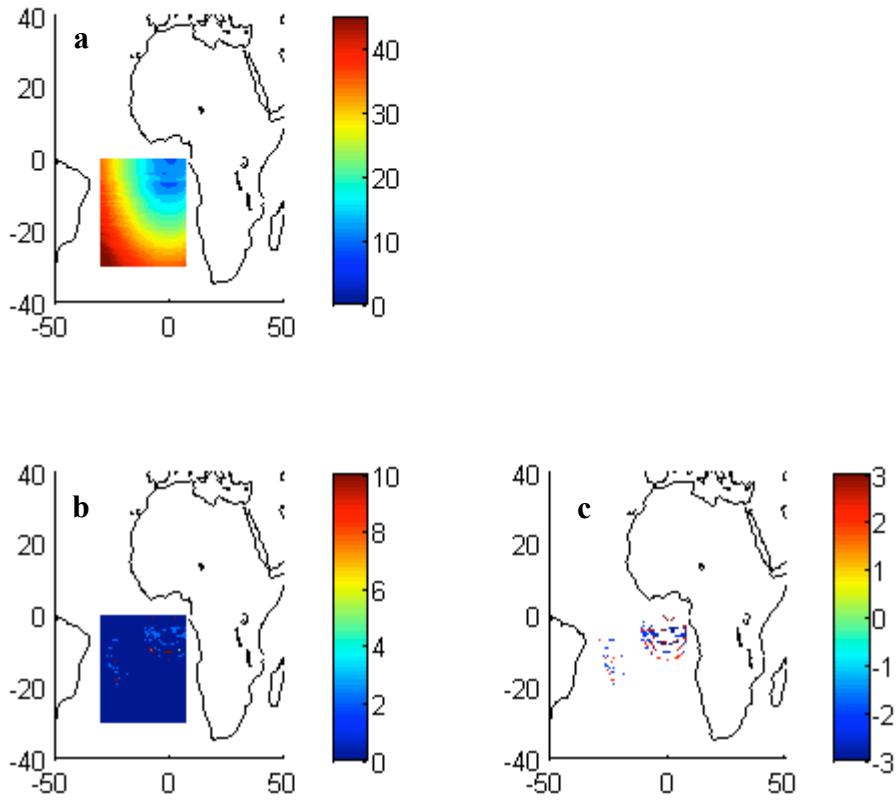


Figure A.1: Satellite viewing angle  $\theta_v$  (a) example of values in January, (b) percentage of values that are more than  $2^\circ$  lower or larger than the median value for the entire time period, (c) mean value of the difference between  $\theta_v$  and its median when the percentage of (b) is non-zero.

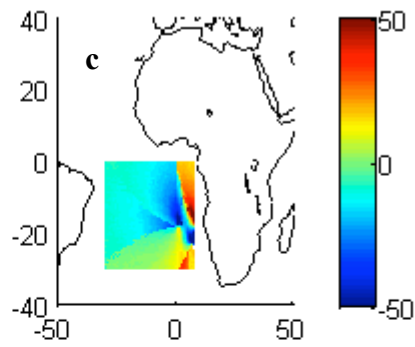
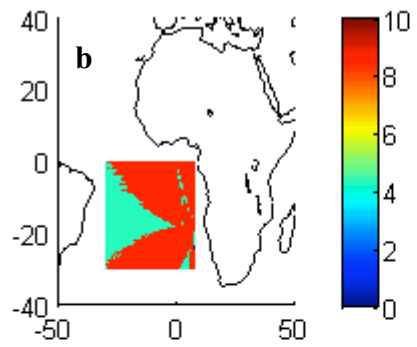
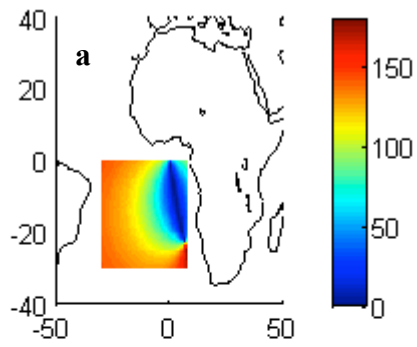
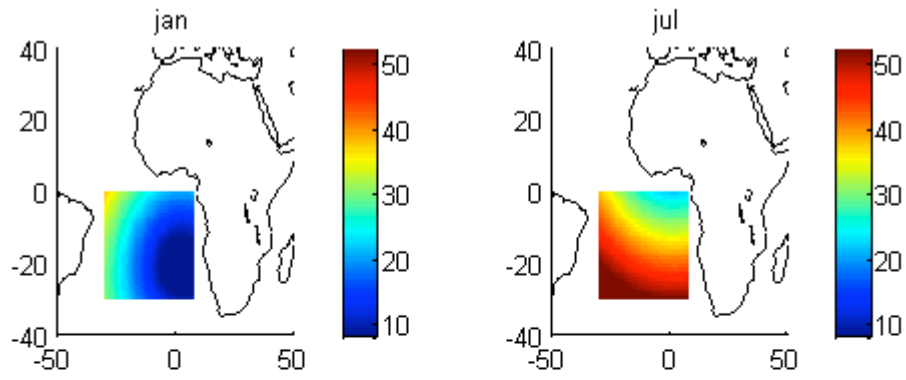


Figure A.2: Satellite azimuthal relative angle  $\phi$  (a) example of values in January, (b) percentage of values that are more than  $5^\circ$  lower or larger than the median value for the entire time period, (c) mean value of the difference between  $\phi$  and its median when the percentage of (b) is non-zero.





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 748 *Figure A.3: (a) Solar zenithal angle at 1200 UTC on January 15<sup>th</sup>; (b) same as (a) but on July*  
 749 *15<sup>th</sup>.*

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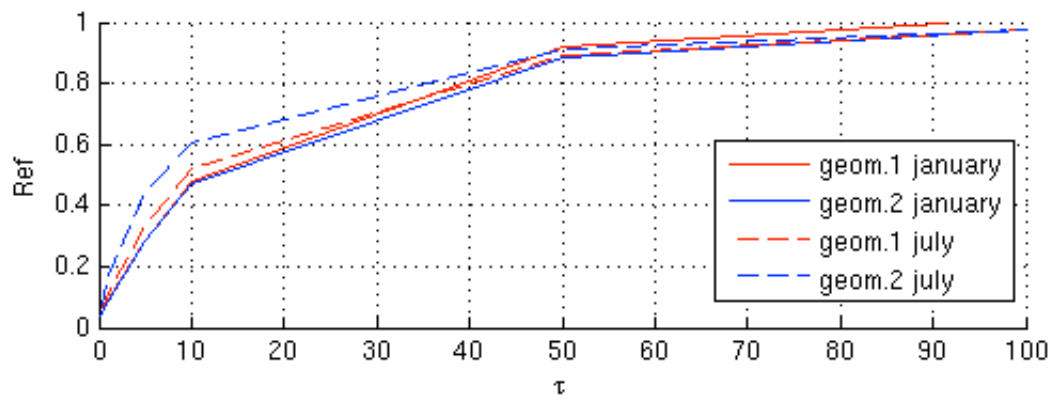
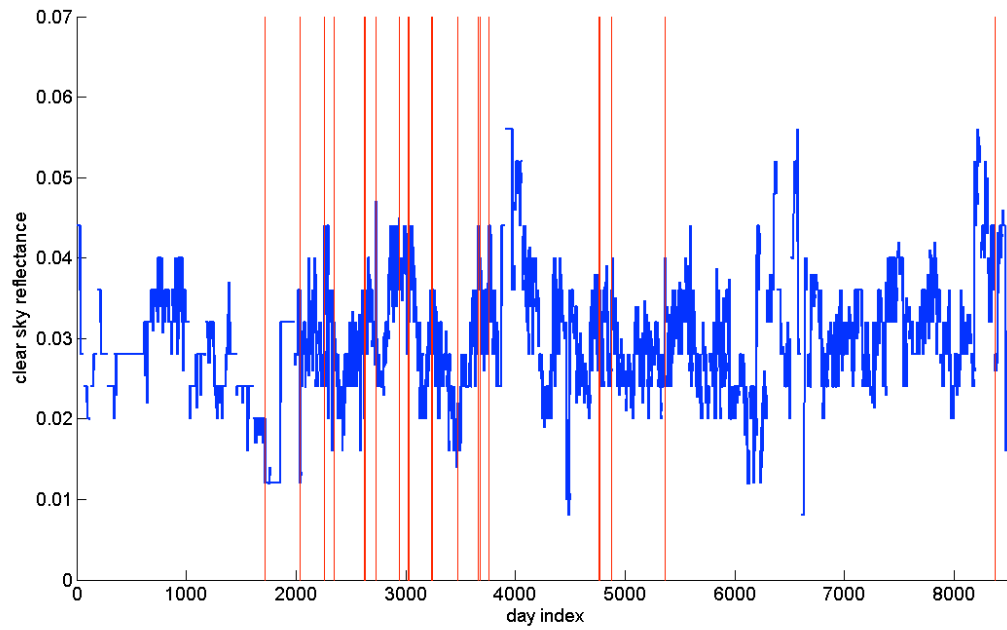


Figure A.4: Simulation of the reflectance as a function of the optical thickness, for four representation of the satellite geometry: solar zenithal angle of  $20^\circ$ /angle of viewing direction with nadir of  $15^\circ$ /relative viewing azimuth angle of  $50^\circ$  (characteristic of January, red plane line),  $32^\circ/15^\circ/50^\circ$  (characteristic of January, blue plane line);  $22^\circ/42^\circ/130^\circ$  (characteristic of July, red dashed line);  $50^\circ/42^\circ/130^\circ$  (characteristic of July, red dashed line).



*Figure A.5: Time evolution of the clear sky reflectance, during the complete period (in blue). Satellite changes are indicated by vertical red lines.*